

X39a ALMA imaging survey for $z \gtrsim 4-5$ 1100- μm -selected galaxy candidates

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We present the initial result of ALMA 1100- μm imaging of 30 Herschel/VLA-faint AzTEC sources in the SXDF/UDS carried out in ALMA cycle-1. We selected 30 AzTEC sources among 281 AzTEC sources by using deep VLA 21 cm, and Herschel 100-500- μm data, aiming to reveal the nature of these ‘submm-radio-faint’ 1100- μm -selected galaxies which are expected to be dusty starburst galaxies at $z \gtrsim 4$. The resultant ALMA images achieve a sensitivity of 70-88 $\mu\text{Jy}/\text{beam}$, $\gtrsim 6$ times better than that of the AzTEC/ASTE observation, and achieve a synthesized beam size of $0''.4-0''.7$ (FWHM), $\gtrsim 2000$ times better than that of AzTEC/ASTE map. We find 35 ALMA continuum sources ($\geq 5\sigma$; $L_{\text{IR}} = 0.4-3.5 \times 10^{12} L_{\odot}$) in total, and 25/30 (=83%) of the ‘submm-radio-faint’ AzTEC sources have at least one significant ALMA continuum source. Deep Optical /Near and Mid infrared data indicate that at least 13/35 (=37%) can be located at $z \gtrsim 4-5$. The baseline lengths of our ALMA data range up to 1200 k λ for 10 AzTEC sources (and up to 400 k λ for remaining 20 sources). We find that all of the 16 bright ALMA/AzTEC sources with ≥ 1 mJy and $\geq 10 \sigma$ detection are resolved, and the median size is ($0''.2$, FWHM). These can be progenitors of the most massive galaxies in the local Universe.