

X33a SXDF-ALMA 1.5 arcmin² deep survey. A compact dusty star-forming galaxy at z=2.5.

但木謙一 (MPE)、河野孝太郎 (東京大学)、児玉忠恭 (国立天文台)、五十嵐創 (Kapteyn
Astronomical Institute, University of Groningen) SXDF-ALMA deep survey チーム

We present results from the SXDF-ALMA 1.5 arcmin² deep survey at 1.1 mm using Atacama Large Millimeter Array (ALMA). The map reaches a 3σ depth of $165 \mu\text{Jy}/\text{beam}$ and covers 11 H α -selected star-forming galaxies at $z = 2.53$ and one at $z = 2.19$. We have detected continuum emission from four of our H α -selected sample. They are all red and have stellar masses of $\log(M_*/M_\odot) > 10.9$ while the other blue, main-sequence galaxies with $\log(M_*/M_\odot)=10.0-10.8$ are exceedingly faint, $< 280 \mu\text{Jy}$ (2σ upper limit), corresponding to a gas mass of $\log(M_{\text{gas}}/M_\odot) = 10.2$ at $z \sim 2.5$. The non-detection raises the possibility that a conversion factor from monochromatic submillimeter luminosity to gas mass is not universal.

The 1.1 mm-brightest galaxy is associated with a compact ($R_e < 1 \text{ kpc}$), dusty star-forming component. Given high gas fraction ($\sim 40\%$) and high star formation rate surface density ($250 M_\odot \text{ yr}^{-1} \text{ kpc}^{-2}$), the centrally-concentrated starburst can within less than 80 Myr build up a stellar surface density within the core matching that of compact quiescent galaxies at $z \sim 2$, provided 20% of the total gas is converted into stars.