

X17a ALMA twenty-Six Arcmin² survey of GOODS-S At One-millimeter (ASAGAO): X-ray AGN Properties of Millimeter-Selected Galaxies

Y. Ueda (Kyoto Univ.), B. Hatsukade, K. Kohno, Y. Yamaguchi (Univ. of Tokyo) , Y. Tamura (Nagoya Univ.), H. Umehata (Open Univ. of Japan), and ASAGAO team

We investigate the X-ray active galactic nucleus (AGN) properties of millimeter galaxies in the GOODS-S field detected with ALMA, by utilizing the Chandra 7-Ms data, the deepest X-ray survey to date. Our millimeter galaxy sample comes from the ASAGAO survey (12 sources at a 1.2-mm flux-density limit of ≈ 0.6 mJy), supplemented by the deeper but narrower 1.3-mm survey of a part of the ASAGAO field by Dunlop et al. (2017). Fourteen out of the total 25 millimeter galaxies have Chandra counterparts. The observed AGN fractions at $z = 1.5 - 3$ is found to be $90_{-19}^{+8}\%$ and $57_{-25}^{+23}\%$ for the ultra/luminous infrared galaxies with $\log L_{\text{IR}}/L_{\odot} = 12-12.8$ and $\log \log L_{\text{IR}}/L_{\odot} = 11.5-12$, respectively. The majority ($\sim 2/3$) of the ALMA-Chandra objects appear to be star-formation dominant populations, having $L_{\text{X}}/L_{\text{IR}}$ ratios smaller than the “simultaneous evolution” value expected from the local black-hole mass to stellar mass ($M_{\text{BH}}-M_{\star}$) relation. On the basis of the L_{X} and stellar mass relation, we infer that a large fraction of star-forming galaxies at $z = 1.5-3$ have black hole masses smaller than those expected from the local $M_{\text{BH}}-M_{\star}$ relation. Our results are consistent with an evolutionary scenario that star formation occurs first, and an AGN-dominant phase follows later, in objects finally evolving into galaxies with classical bulges.