

## X48a Discovery of an Extremely Luminous Dust-obscured Galaxy Observed with SDSS, WISE, JCMT, and SMA

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We present the discovery of an extremely luminous dust-obscured galaxy (DOG) at  $z_{\text{spec}} = 3.703$ , WISE J101326.25+611220.1. This DOG is selected as a candidate of extremely luminous infrared (IR) galaxies based on the photometry from the SDSS and WISE. In order to derive its accurate IR luminosity, we perform follow-up observations at 450 and 850  $\mu\text{m}$  using the SCUBA2 on the JCMT, and at 870 and 1300  $\mu\text{m}$  using the SMA, which enable us to pin down its IR Spectral Energy Distribution (SED). We perform SED fitting using 14 photometric data (0.4–1300  $\mu\text{m}$ ) and estimate its IR luminosity,  $L_{\text{IR}}$  (8–1000  $\mu\text{m}$ ), to be  $2.2_{-1.0}^{+1.5} \times 10^{14} L_{\odot}$ , making it one of the most luminous IR galaxies in the universe. The energy contribution from an active galactic nucleus (AGN) to the IR luminosity is  $94_{-20}^{+6}\%$ , which indicates that it is an AGN-dominated DOG. On the other hand, its stellar mass ( $M_*$ ) and star formation rate (SFR) are  $\log(M_*/M_{\odot}) = 11.2_{-0.2}^{+0.6}$  and  $\log(\text{SFR}/M_{\odot} \text{ yr}^{-1}) = 3.1_{-0.1}^{+0.2}$ , respectively, which means that this DOG can be considered a starburst galaxy in the  $M_*$ –SFR plane. This extremely luminous DOG shows significant AGN and star-forming activity that provides us with an important laboratory to probe the maximum phase of the coevolution of galaxies and supermassive black holes (Toba et al. 2018, ApJ, 857, 31).