

S07a A robust lower bound on intergalactic magnetic fields from Fermi/LAT and MAGIC observations of 1ES 0229+200

I. Vovk (ICRR, U. Tokyo), A. Stamerra (INAF), P. Da Vela (INFN, Innsbruck U.), A. Neronov (U. Paris, EPFL), D. Semikoz (U. Paris, INR RAS, MEPhI), A. Korochkin (U. Paris, INR RAS, Novosibirsk U.), K. Asano, G. Ceribella, D. Hadasch, M. Hütten, T. Inada, Y. Iwamura, Y. Kobayashi, K. Noda, Y. Ohtani, T. Saito, S. Sakurai, J. Sitarek, M. Strzys, M. Takahashi, R. Takeishi (ICRR), D. Mazin, M. Teshima (ICRR, MPI Phys.), Y. Fukazawa, Y. Suda (Hiroshima U.), H. Kubo, S. Nozaki, T. Oka (Kyoto U.), J. Kushida, K. Nishijima (Tokai U.), T. Nakamori (Yamagata U.), T. Yamamoto (Konan U.) on behalf of the MAGIC Collaboration

Secondary  $\gamma$ -ray emission from distant TeV sources induced by the effects of propagation of  $\gamma$  rays through the intergalactic medium could be used to measure the intergalactic magnetic field (IGMF). A proper realization of this opportunity requires a knowledge of the past source TeV luminosity evolution over the relevant period of time. Here we use the sample of MAGIC, H.E.S.S., VERITAS and Fermi/LAT observations to trace evolution of the hard-spectrum blazar 1ES0229+200 in the GeV-TeV band over  $\sim 15$  years. This allows us to make a precise prediction of the timing properties of the time-delayed secondary  $\gamma$ -ray flux. We show that the non-detection of such an emission in GeV energy band yields a robust lower bound on the strength of IGMF.