

X37a Properties of low-mass starburst $H\alpha$ emitters (Flash-HAEs) at $z \sim 2.3$ from broad-band selection

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Recent studies reveal that low-mass galaxies at $z > 6$ are expected to be major contributors to cosmic reionization. Because of the observational limitation, a practical route to grasp the physical properties of these early galaxies is to undertake a detailed study on their lower-redshift analogs. We investigate the properties of ~ 1800 $H\alpha$ emitters (HAEs) at $z \sim 2.3$, selected from the flux excess in the ZFOURGE Ks broad-band data to the best-fit stellar continuum. A new population called low-mass starburst $H\alpha$ emitters (Flash-HAEs) is identified from their elevated $SFR(H\alpha)$ above the star formation main sequence (SFMS) in low-mass regime below $10^9 M_\odot$. The 306 Flash-HAEs in our work may be an analog population of the low-mass galaxies at the epoch of reionization. The hydrogen ionizing photon production efficiencies, ξ_{ion} , of the Flash-HAEs have a median value of $\log(\xi_{ion}/erg^{-1} Hz) = 25.18 \pm 0.16$, higher than that of the main sequence galaxies by ~ 0.2 dex at similar redshift. This indicates that the Flash-HAEs are more efficient in producing ionizing photons. We will further discuss physical properties of the Flash-HAEs with multiple rest-frame optical emission lines obtained from medium band filters. These new findings would provide implication to the low-mass galaxies at the epoch of reionization.