

X56a RIOJA: Optical and Far-Infrared View of a Major Merger at $z = 7.15$

Y. Sugahara^{1,2}, J. Álvarez-Márquez³, T. Hashimoto⁴, L. Colina³, A. K. Inoue¹, L. Costantin³, Y. Fudamoto^{1,2}, K. Mawatari⁴, S. Arribas³, T. J. L. C. Bakx⁵, C. Blanco-Prieto³, D. Ceverino⁶, A. Crespo Gómez³, M. Hagimoto⁵, T. Hashigaya⁷, R. Marques-Chaves⁸, H. Matsuo², Y. Nakazato⁹, M. Pereira-Santaella¹⁰, Y. W. Ren¹, Y. Tamura⁵, M. Usui⁴, N. Yoshida⁹ (¹Waseda U., ²NAOJ, ³El Centro de Astrobiología, ⁴Tsukuba U., ⁵Nagoya U., ⁶U. Autònoma de Madrid, ⁷Kyoto U., ⁸Geneva Observatory, ⁹U. of Tokyo, ¹⁰Instituto de Física Fundamental)

Combinations of JWST infrared and ALMA submillimeter observations allow us to investigate a multi-phase structure of high-redshift galaxies. The Reionization and the ISM/Stellar Origins with JWST and ALMA (RIOJA) is a JWST GO Cycle 1 program targeting bright [OIII]88+[CII]158 μm emitters at $z > 6$ with NIRCcam and NIRSpec IFU to establish relations between stellar, gaseous, and dusty properties of the high-redshift galaxies. In this talk, we would like to present results of our NIRCcam observations targeting B14-65666 (a.k.a. Big Three Dragons), a bright major merger at $z = 7.15$. The NIRCcam images clearly revealed detailed morphology of the two merging clumps in mid-infrared wavelengths in the observed frame. A merging clump is accompanied with diffuse emission around it, indicating large tidal tails at this high redshift. The F444W photometry shows excesses in both merging clumps being caused by a strong [OIII] 5007 emission line, which support that it is experiencing instant starburst caused by the major merging event. We will discuss results of the SED fitting and a comparison with ALMA observations of emission lines and underlying dust continua.