

P310b An improved upper limit of Cyanodiacetylene ( $\text{HC}_5\text{N}$ ) in the atmosphere of Titan

Takahiro Iino (Univ. of Yamanashi, The Univ. of Tokyo), Kotomi Taniguchi (NAOJ), Hideo Sagawa (Kyoto Sangyo Univ., The Univ. of Tokyo), Martin. A. Cordiner (NASA GSFC, Catholic Univ. of America), Shuya Tan (JAMSTEC), Shigeru Takahashi (Univ. of Hyogo) and Hikaru Kubota (Kyoto Sangyo Univ.)

In the complex Titan atmospheric chemistry, photochemical reactions involving cyanodiacetylene ( $\text{HC}_5\text{N}$ ) may be important as reaction pathways that could lead to the production of amines and aromatic molecules. We carried out a new comprehensive search of  $\text{HC}_5\text{N}$  in Titan's atmosphere using Band 6 spectroscopic archival data from the ALMA. To search for a faint spectral line feature of  $\text{HC}_5\text{N}$ , we used a spectral stacking method that integrates different transition data. Integrating six independent observation datasets with a total integration time of  $\sim 15000$  seconds, we found no statistically significant detection of an  $\text{HC}_5\text{N}$  spectral feature. Using the radiative transfer modeling and employing the four previously predicted vertical distributions, we derived the upper limit column density of  $\text{HC}_5\text{N}$ , corresponding to  $3\text{-}\sigma$  noise level, to be  $0.3 - 10 \times 10^{13}$  molecules/cm<sup>2</sup> above  $\sim 60$  km altitude. The corresponding upper limit of  $[\text{HC}_5\text{N}]/[\text{HC}_3\text{N}]$  ratio value derived from three of four models seem to be lower than that measured in High-mass Protostellar Objects, possibly due to the lower abundance of  $\text{C}_4\text{H}_2$ . This work was published in the *Astronomical Journal* (DOI: 10.3847/1538-3881/ae0621).